

Laser-pulse-shape control of seeded QED cascades

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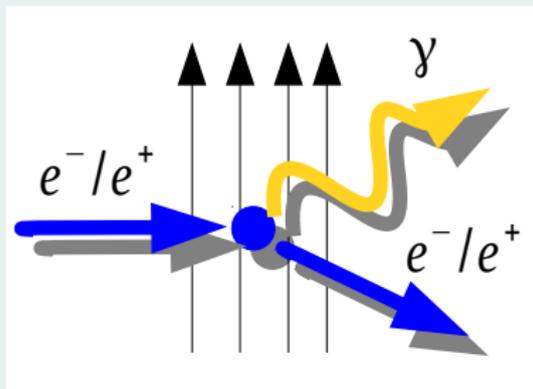
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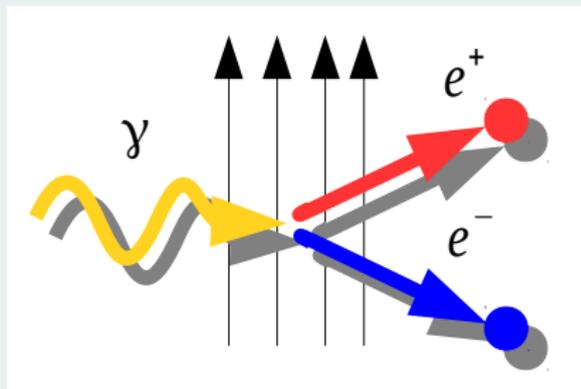
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The basic QED processes in a strong background electromagnetic field

Photon emission by an e^-/e^+



Photon conversion into an e^-e^+ pair



Strong-field QED processes are controlled by the quantum parameter χ

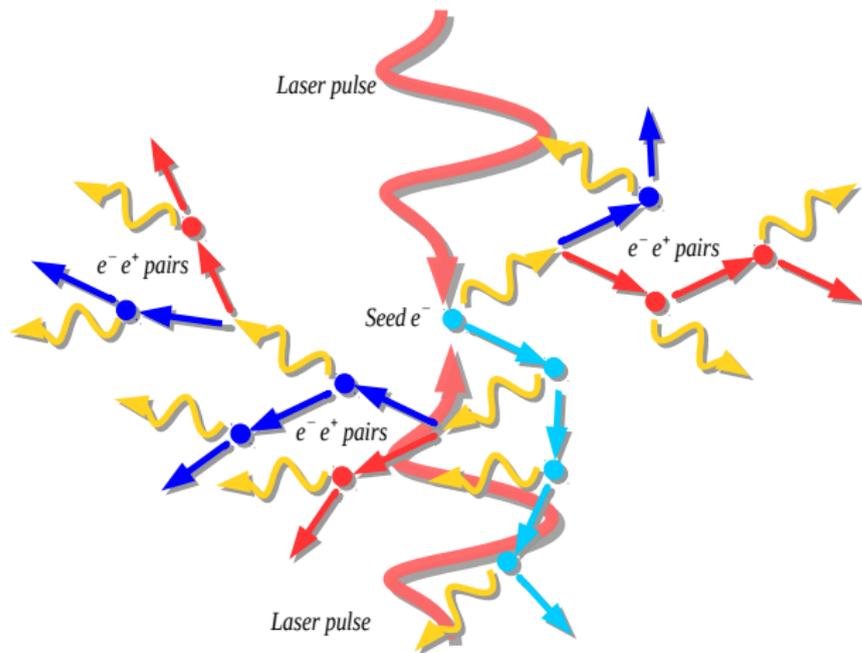
$$\chi_{e/\gamma} = |F_{\mu\nu} p_{e/\gamma}^\nu| / E_{cr} mc$$

For $\chi_e \ll 1$ the typical energy of the emitted photons is $\varepsilon_\gamma \approx \chi_e \varepsilon_e$, ε_e being the e^-/e^+ energy. **Single photon emission recoil dominates when $\chi_e \gtrsim 1$.**

$$eE_{cr}\lambda_C = mc^2; \quad E_{cr} \approx 1.3 \times 10^{16} \text{ V/cm}$$

For $\chi_\gamma \ll 1$ the probability of photon conversion into an e^-e^+ pair is suppressed as $e^{-8/3\chi_\gamma}$. **Photon conversion is important when $\chi_\gamma \gtrsim 1$.**

Laser-driven seeded QED cascades



1) Seed e^- are violently accelerated by the laser fields and emit large amounts of γ which, in turn, convert into $e^- e^+$ pairs.

2) The generated $e^- e^+$ pairs are then accelerated by the laser fields and originate a new generation of particles.

3) QED cascades were predicted to develop in the collision of two laser pulses, each pulse with an intensity around 10^{24} W/cm^2 (Bell *et al.* PRL 2008, Fedotov *et al.* PRL 2010, Nerush *et al.* PRL 2011).

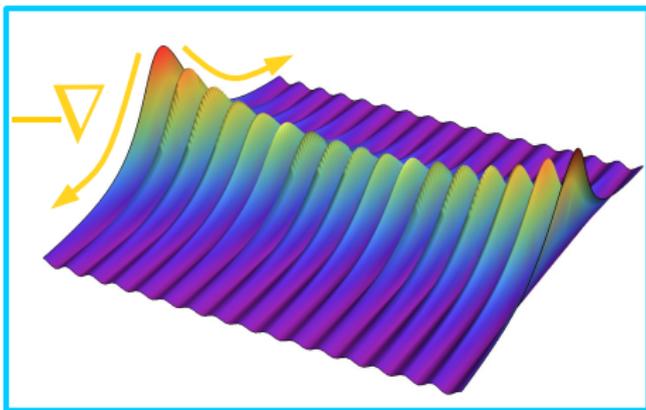
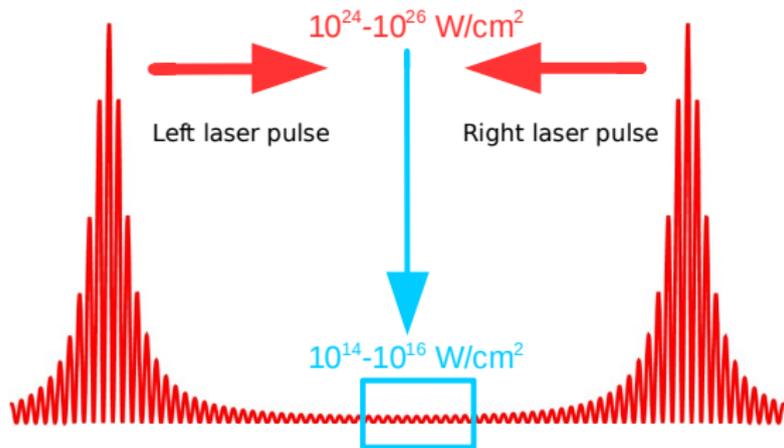


Why are laser-driven QED cascades interesting?

- QED cascades open up the investigation of a **new regime dominated by the interplay between strong-field QED and multiparticle processes**. Raising number of publications! (Bell *et al.* PRL 2008, Fedotov *et al.* PRL 2010, Nerush *et al.* PRL 2011, Bulanov *et al.* PRA 2013, Gelfer *et al.* PRA 2015, Jirka *et al.* PRE 2016, Gonoskov *et al.* arXiv 2016, Zhu *et al.* NC 2016, Vranic *et al.* PPCF 2016, Grismayer *et al.* PRE 2017).
- QED cascades play a **fundamental role in astrophysical environments** such as the magnetosphere of pulsars, rendering an earth based implementation with intense lasers attractive (**relativistic laboratory astrophysics** viable with $\gtrsim 10$ PW lasers).
- QED cascades can limit the attainable intensity of extreme laser sources due to the **depletion of the laser pulse energy** (Fedotov *et al.* PRL 2010, Nerush *et al.* PRL 2011).

Our goal: how to control the onset of seeded QED cascades?

- Hitherto, the research has been **focused on the intensity** required to trigger QED cascades. The implications of the strong field gradients associated with tightly focused laser pulses have been neglected.
- We have shown (Tamburini *et al.*, Sci. Rep. **7**, 5694 (2017)) the **essential role played by the laser field structure** (e.g. waist radius and duration, relative orientation of the polarization axis) on the onset of seeded QED cascades.



Field gradients

The presence of **field gradients** results into the **expulsion of e^-e^+** from the regions where the field is stronger.

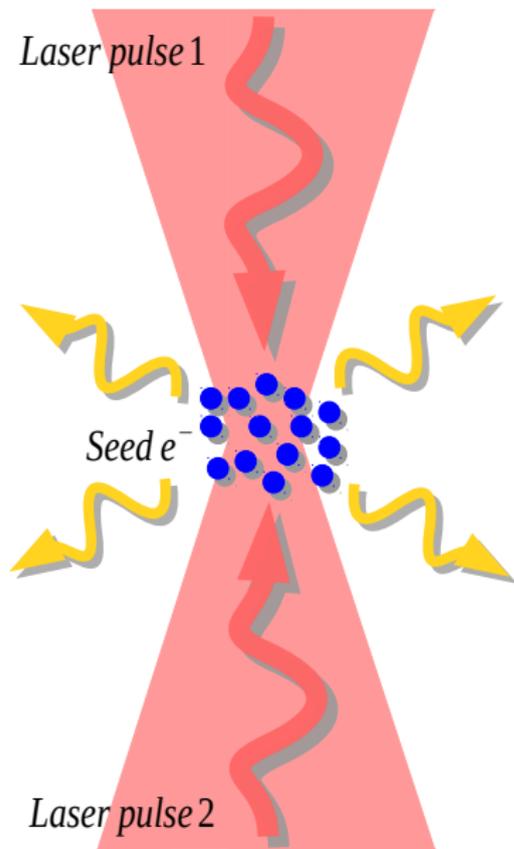
For instance, for a single laser pulse described by the vector potential $\mathbf{A}(\mathbf{r}, t)$, the secular e^-e^+ dynamics is known to be determined by the **ponderomotive force**:

$$\mathbf{f}_p = -mc^2 \nabla (1 + \langle \mathbf{a}^2 \rangle)^{1/2}$$

where $\mathbf{a} = e\mathbf{A}/mc^2$ is the normalized vector potential.

To attain the **highest intensities** \Rightarrow **tight focusing** \Rightarrow **strong field gradients**.

Thus, the expulsion of e^-e^+ may play a **critical role** on the onset of QED cascades.

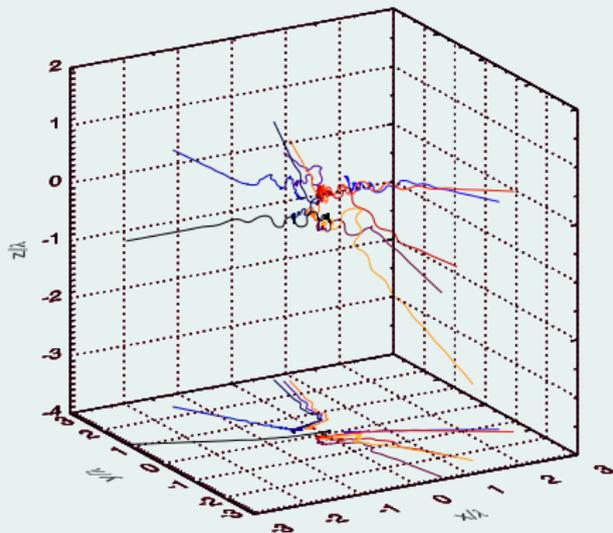


Parameters

- Two ultraintense **linearly polarized** (LP_{\parallel} or LP_{\perp}) laser pulses collide head-on in a tenuous gas (such as the residual gas of a vacuum chamber).
- Their transverse spatial profile is Gaussian, with $\lambda = 0.8 \mu\text{m}$ wavelength, and sech^2 intensity envelope with 20 fs duration FWHM. Initially, the peaks of the two laser pulses are located at $z_0 = \pm 70 \lambda$.
- A fully 3D description of the laser pulse fields with terms up to the fifth order in the diffraction angle $\epsilon = \lambda/\pi w_0$, w_0 being the waist radius, is employed.
- Initially, 10^3 seed electrons are located at rest within a λ^3 volume at the laser pulse focus with uniform random distribution. The **electron density is $n = 2 \times 10^{15} \text{cm}^{-3} \approx 10^{-6} n_c$** , where the critical density is $n_c = m\omega^2/4\pi e \approx 1.7 \times 10^{21} \text{cm}^{-3}$.
- Seed electrons originate from the ionization of different atomic species (e.g. H, O), and **go into the continuum at different values of the laser field at the focus.**

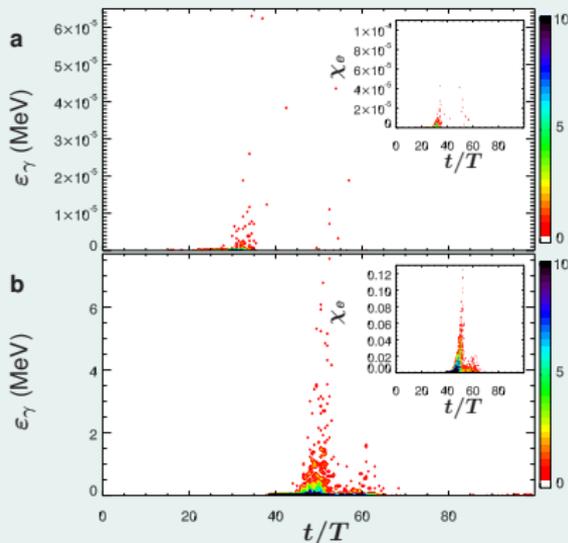


Trajectories of ten seed electrons



The trajectories for $0 < t < 36.5 T$ of ten seed electrons. At $t = 36.5 T$ all seed electrons are outside the focal spot and the laser pulse peaks are still at $z = \pm 33.5 \lambda$. The projection of the trajectories on the focal plane xy is also reported.

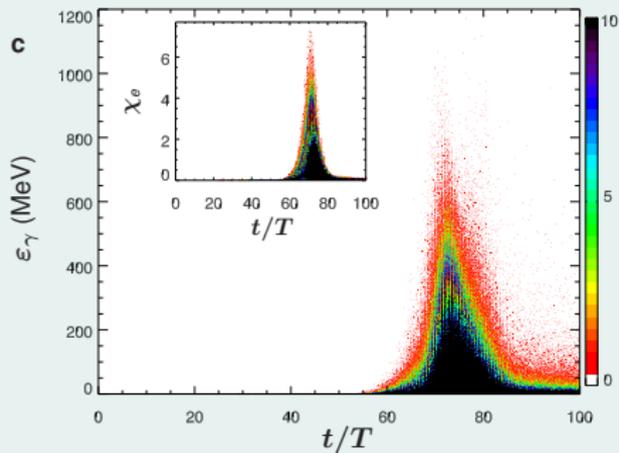
Emitted photon energies



The photon energy ϵ_γ and the χ_e parameter (inset) at each photon emission event as function of time. Initially, the peaks of the two laser pulses are located at $z_0 = \pm 70 \lambda$. (a) Hydrogen (b) Hydrogen-like Oxygen. The colors correspond to the number of events (black means ≥ 10 events).

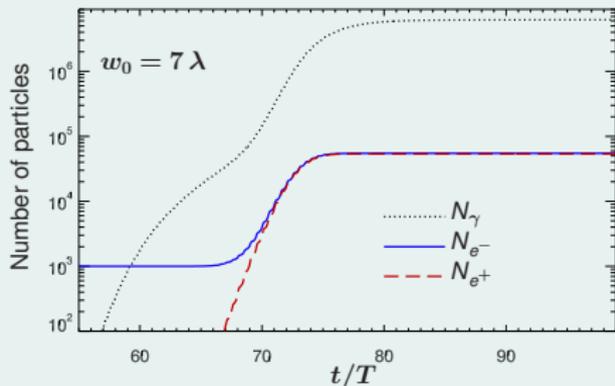


Emitted photon energies



The photon energy ε_γ and the χ_e parameter (inset) at each photon emission event. Initially, the peaks of the laser pulses are located at $z_0 = \pm 70 \lambda$ and electrons originate from hydrogen. The colors correspond to the number of events (black means ≥ 10 events).

The number of particles



The evolution of the number of electrons N_{e-} , positrons N_{e+} and photons N_γ with energy $\varepsilon_\gamma > 2.5$ MeV. Although the laser pulse intensity decreases from 4.8×10^{25} W/cm² to 10^{24} W/cm² here χ_e exceeds unity and **copious emission of photons with hundreds MeV energy occurs.**

The onset of QED cascades as function of the laser pulse power P and waist radius w_0 , both for LP_{\parallel} and LP_{\perp} .

P (PW)	w_0 (λ)	LP_{\parallel} (H O)		LP_{\perp} (H O)	
200	1	N	N	N	N
	2	N	$C_{2.4}$	N	N
	3	N	$C_{1.4}$	N	$N \leftrightarrow C$
	4	$G \leftrightarrow C$	$C_{0.7}$	N	$G \leftrightarrow C$
	5	G	$C_{0.3}$	G	G
	[6, 9]	G	G	G	G
500	1	N	N	N	N
	2	N	$C_{6.0}$	N	N
	3	$N \leftrightarrow C$	$C_{3.7}$	N	$N \leftrightarrow C$
	4	$C_{1.6}$	$C_{2.4}$	N	$C_{2.1}$
	[5, 8]	$C_{[1.3, 0.3]}$	$C_{[1.7, 0.4]}$	$C_{[0.7, 0.3]}$	$C_{[1.1, 0.2]}$
	9	$C_{0.2}$	$C_{0.2}$	$G \leftrightarrow C$	$G \leftrightarrow C$
	[10, 15]	G	G	G	G

N = No e^-e^+ pairs, G = e^-e^+ gas, C = $e^-e^+\gamma$ cascade, $A \leftrightarrow B$ = transition between regime A and B. The subscripts indicate the average cascade growth rate $\langle \Gamma \rangle$ in units of T^{-1} .

The onset of QED cascades as function of the laser pulse intensity I and waist radius w_0 , both for LP_{\parallel} and LP_{\perp} .

I (W/cm^2)	w_0 (λ)	LP_{\parallel} (H O)		LP_{\perp} (H O)	
10^{24}	1	N	N	N	N
	2	N	G	N	N
	3	N \leftrightarrow G	G \leftrightarrow C	N	N \leftrightarrow G
	4	N \leftrightarrow G	C	N \leftrightarrow G	G
	5	G \leftrightarrow C	C	G	G
	≥ 6	C	C	C	C
10^{25} & 10^{26}	1	N	N	N	N
	2	N	C	N	N
	3	N \leftrightarrow C	C	N	N \leftrightarrow C
	4	C	C	N \leftrightarrow C	C
	≥ 5	C	C	C	C

N = No e^-e^+ pairs, G = e^-e^+ gas, C = $e^-e^+\gamma$ cascade, A \leftrightarrow B = transition between regime A and B.

Our main findings on the onset of seeded QED cascades

- The laser field structure is of critical importance for the onset of e^-e^+ cascades. The laser intensity is not the only decisive parameter.
- We have highlighted the importance of the nature of the gas. Inner shell electrons of high- Z elements may go into the continuum only when the peaks of the laser pulses reach the focus. In this case the power required to initiate a QED cascade falls to 11 PW per pulse.
- The laser pulse structure is of critical importance also in more complex setups with, e.g., four colliding laser pulses.
- These findings open up the possibility of controlling the onset of seeded QED cascades. Seeded QED cascades can be:
 - 1) facilitated with $w_0 \gtrsim 4\lambda$ and high- Z gases at $I \lesssim 10^{24} \text{ W/cm}^2$
 \Rightarrow relativistic laboratory astrophysics,
 - 2) prevented with $w_0 \lesssim 2\lambda$ and low- Z gases up to $I \approx 10^{26} \text{ W/cm}^2$
 \Rightarrow attain the highest intensities and probe the QED vacuum.